

# Lecture 14: Equilibrium Equations

Logistics: - HW 5 due 3/7 ! ! !  
- HW 6 will be posted

Last time: - curl  $(\nabla \times \underline{v}) \times \underline{a} = (\nabla \underline{v} - \nabla \underline{v}^T) \underline{a}$

$$\underline{\omega} = \nabla \times \underline{v} = \epsilon_{ijk} v_{i,k} \underline{e}_j$$

$$\nabla \times \nabla \phi = 0 \quad \nabla \cdot (\nabla \times \underline{v}) = 0$$

- Laplacian:  $\nabla \cdot \nabla = \Delta = \nabla^2$

scalar  $\Delta \phi = \phi_{,ii}$

vector  $\Delta \underline{v} = v_{i,jj} \underline{e}_i$

$$\Delta \underline{v} = \nabla(\nabla \cdot \underline{v}) - \nabla \times \nabla \times \underline{v}$$

- Integral laws

Div. Theorem:  $\int_{\partial \Omega} \underline{v} \cdot \underline{n} \, dS = \int_{\Omega} \nabla \cdot \underline{v} \, dV$

$$\int_{\partial \Omega} \underline{s} \cdot \underline{n} \, dS = \int_{\Omega} \nabla \cdot \underline{s} \, dV$$

Stokes Thm:  $\int_{\Gamma} (\nabla \times \underline{v}) \cdot \underline{n} \, dA = \oint_{\partial \Gamma} \underline{v} \cdot \underline{t} \, ds$

$$|\nabla \times \underline{v}| = \omega$$

Today: - Equilibrium equations

- Hydrostatic shapes

# Mechanical Equilibrium

Necessary conditions for eqbm

$$\underline{\Gamma}[\Omega] = \underline{\Gamma}_b[\Omega] + \underline{\Gamma}_s[\partial\Omega] = \int_{\Omega} \rho \underline{b} dV + \int \underline{t} dA = 0$$

$$\underline{C}[\Omega] = \underline{C}_b[\Omega] + \underline{C}_s[\partial\Omega] = \int_{\Omega} \underline{x} \times \rho \underline{b} dV + \int_{\partial\Omega} \underline{x} \times \underline{t} dA = 0$$

If  $\underline{\Gamma}[\Omega] = 0$  then  $\underline{C}[\Omega]$  is independent of  $\underline{z}$ !

Substitute definition of Cauchy stress  $\underline{t} = \underline{\underline{\sigma}} \underline{n}$   
 $\underline{n}$  = unit normal

$$\underline{\Gamma}[\Omega] = \int_{\partial\Omega} \underline{\underline{\sigma}} \underline{n} dA + \int \rho \underline{b} dV = 0$$

Tensor divergence theorem

$$\underline{\Gamma}[\Omega] = \int_{\Omega} \nabla \cdot \underline{\underline{\sigma}} + \rho \underline{b} dV = 0$$

by arbitrariness of  $\Omega$  we have

$$\underline{\nabla \cdot \underline{\underline{\sigma}} + \rho \underline{b}} = 0$$

$$\underline{C}[\Omega] = \int_{\partial\Omega} \underline{x} \times (\underline{\underline{\sigma}} \underline{n}) dA + \int_{\Omega} \underline{x} \times \rho \underline{b} dV$$



$$\epsilon_{ijk} \delta_{jk} \sigma_{kl} = \boxed{\epsilon_{ijk} \sigma_{kj} = 0}$$

If  $\epsilon_{ijk} \sigma_{kj} = 0$  then  $\epsilon_{ikj} \sigma_{jk} = 0$   
because  $j, k$  are dummy indices

$$0 = \epsilon_{ijk} \sigma_{kj} + \epsilon_{ikj} \sigma_{jk} = \epsilon_{ijk} (\sigma_{kj} - \sigma_{jk}) = 0$$

Can choose  $i$  different from  $j$  &  $k$   $\epsilon_{ijk} \neq 0$

$$\Rightarrow \boxed{\sigma_{kj} = \sigma_{jk}}$$

$\Rightarrow$  stress tensor is symmetric  $\underline{\underline{\sigma}} = \underline{\underline{\sigma}}^T$

Equations for mechanical Eqm

$$\boxed{\begin{aligned} \nabla \cdot \underline{\underline{\sigma}} + \rho \underline{\underline{b}} &= \underline{\underline{0}} & \sigma_{ij,i} + \rho b_i \\ \underline{\underline{\sigma}}^T &= \underline{\underline{\sigma}} & \sigma_{ij} = \sigma_{ji} \end{aligned}}$$

force bal.

torque bal.



## I) Stationary body

Homogeneous body

$$\Phi = \begin{cases} \frac{4}{3}\pi G\rho (r^2 - R^2) & r \leq R \\ -\frac{1}{3}MG \left(\frac{1}{r^3} - \frac{1}{R^3}\right) & r \geq R \end{cases}$$

$\Phi = \Phi(r) \Rightarrow$  equipotentials are spheres

## II) Rotating body

In frame rotating with body  $\rightarrow$  centrifugal force

$$\underline{f}_c = m\Omega^2 \underline{s}$$

$\Omega =$  angular velocity

$\underline{s} = \begin{pmatrix} x \\ y \end{pmatrix}$  shortest distance  
from rotation axis

$s = |\underline{s}|$  dist. from axis

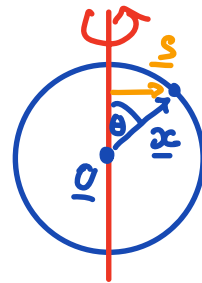
$\theta =$  polar angle

$$\underline{s} = \underline{x} \sin \theta$$

rotation adds "fictitious" accelerations & potential

centrifugal:  $\underline{g}_c = \Omega^2 \underline{s}$       $\Phi_c = -\frac{1}{2}\Omega^2 s^2$       $\underline{g}_c = -\nabla\Phi_c$

gravitational:  $\underline{g}_G = \frac{GM}{R^2}$       $\Phi_G \approx |g| dr = \frac{GM}{R^2} dr$



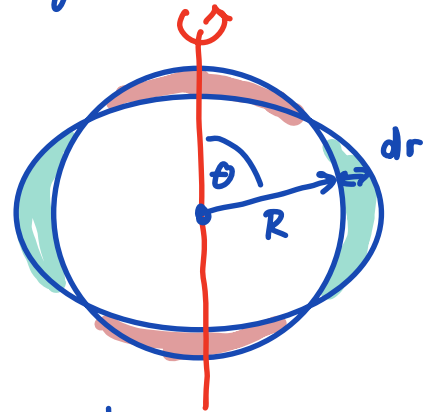
note:  $\bar{\Phi}_G$  is linearized at surface

Total potential:

$$\bar{\Phi} = \bar{\Phi}_G + \bar{\Phi}_C + \Delta\bar{\Phi}$$

$\Delta\bar{\Phi}$  = self potential

change of  $\bar{\Phi}_G$  due to deformation



Ignore  $\Delta\bar{\Phi}$

$$\bar{\Phi} = \frac{MG}{R^2} \underline{dr}(\theta) - \frac{1}{2} \frac{\Omega^2 R^2}{2} \sin^2 \theta + \text{const.}$$

solve for  $dr$

$$dr = \frac{R}{2} q \sin^2 \theta + \text{const}$$

$$q = \frac{\Omega^2 R^3}{MG} = \frac{|g_c|}{|g_s|}$$

Constant determined by mass/volume conservation

$$\int_0^\pi dr(\theta) dS = 2\pi R^2 \int_0^\pi dr(\theta) \sin \theta d\theta = 0$$

$$dr = dr_0 \left( \sin^2 \theta - \frac{2}{3} \right)$$

$$dr_0 = \frac{1}{2} R q$$

Earth:  $R = 6371 \text{ km}$

$g_s = 9.81 \frac{\text{m}^2}{\text{s}^2}$      $\Omega = 7.6 \cdot 10^{-6} \frac{1}{\text{s}}$

$$\Rightarrow q \approx 3.5 \cdot 10^{-3}$$

$$\Rightarrow dr_0 \approx 11 \text{ km}$$

actual  $dr_0 = 21.4 \text{ km}$

Error is due to self-potential  $\Delta\Phi$

$$\Delta\Phi = -\frac{3}{5} \frac{\rho_1}{\rho_0} g_s dr(\theta)$$

where  $\rho_0 = \frac{M}{V}$  and  $\rho_1$  density of displaced material

$$\rho_0 = 5500 \frac{\text{kg}}{\text{m}^3} \quad \rho_1 = 4500 \frac{\text{kg}}{\text{m}^3}$$

$$\Rightarrow dr_0 = \frac{1}{2} R g \left( 1 - \frac{3}{5} \left( \frac{\rho_1}{\rho_0} \right) \right)^{-1} \approx 21.9 \text{ km}$$