

Lecture 26: Wrapping up bits & pieces

Logistics: • all HW's are done ✓ Thank you! ∇

- course evaluations
- Thursday → feedback.

Last time: Lecture 23: Steady coupled ocean-GW model

Lecture 24: Filling craters

⇒ BC is integral constraint

$$h'(1) \sim \int_0^1 h' dx'$$

Lecture 25: Mars topography

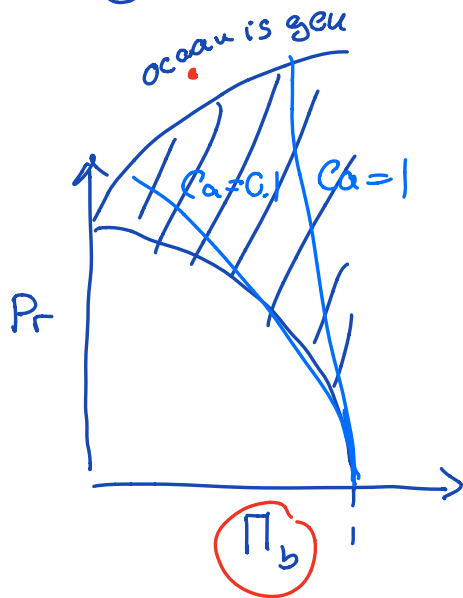
Today: - go over topography live script

- Solve problem with surface water

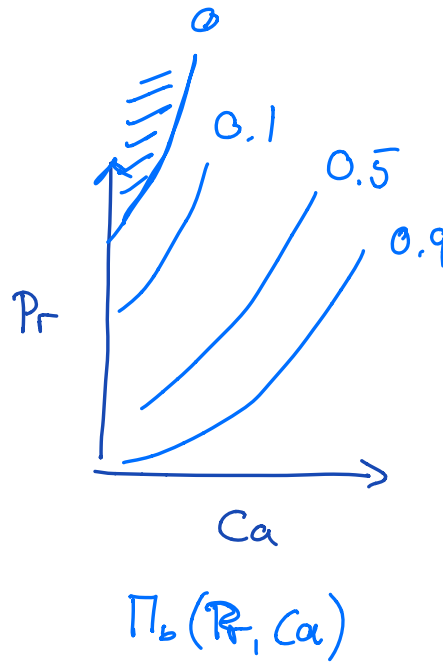
coupling numerically with Newton-Raphson

- Summary of what we have learned

Steady GW problem



head on bud



Lake filling problem

$$\text{PDE: } \phi \frac{\partial h}{\partial t} - K \nabla \cdot u \nabla u = 0$$

$$\text{ODE: } \omega \frac{dh_e}{dt} = (h_e)_0$$

↑

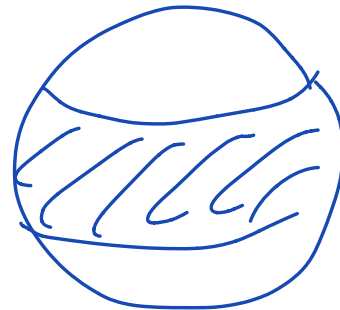
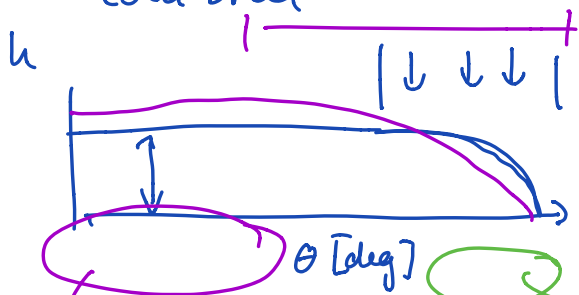
That's it folks! ☹

Outline Mars GW paper

1) Steady solutions for confined & unconfined flow on a spherical shell.

- with precip
- with recharge

• combined



very long residence times (pink) short residence times (green)

⇒ show how head changes with precip distribution

learn that comb. of decay of h with depth + precip near equator makes for a very flat GW with steep decline near the edge

⇒ $h \sim$

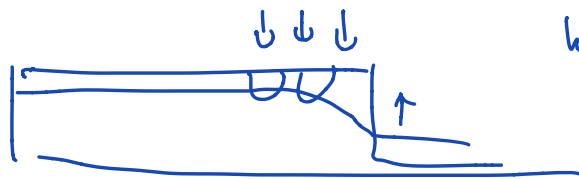
$v \sim$

⇒ residence times

2) Steady coupled GW-ocean problem

→ extended to physical shell & changing overall planet scale water balances

100 - 1500 m global layer

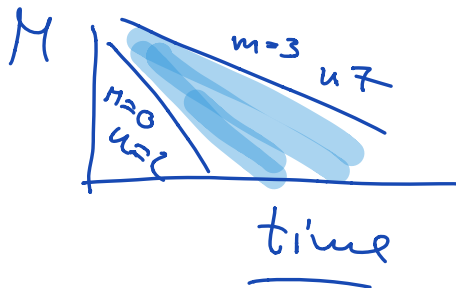
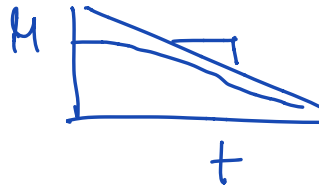
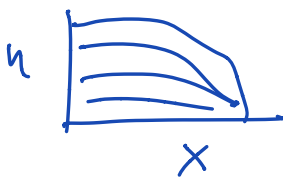


how much precip

do I need to

raise GW to inferred height

3) Transient decay of GW



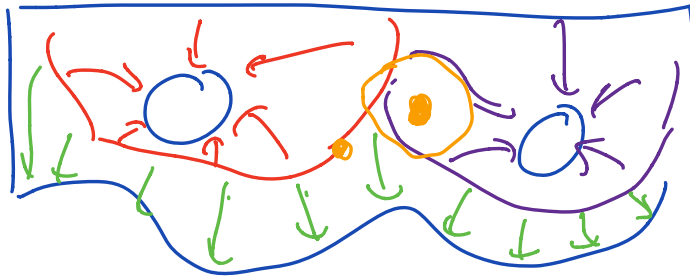
$$\text{PDE: } \frac{\partial h^{m+1}}{\partial t} + D_n \nabla \cdot (h^{n+1} \nabla h) = 0$$

$$\text{ODE: } \frac{m+1}{n-m+1} f^{m+1} + \frac{d}{dx} \left(f^{n+1} \frac{df}{dx} \right) =$$

same structure

$$\frac{m+1}{n-m+1} f^{m+1} + \nabla \circ (f^{m+1} \nabla f) = 0$$

This self similar soln exists on complex geometry



⇒ allow to map global drainage patterns of Rastrian highlands.