

Gravity

In many planetary problems g varies and has to be computed by solving the original Poisson equation for the gravitational field:

Gravitational field: $\mathbf{g} = -\nabla\Phi$ (vector field)
where Φ is gravitational potential

Gauss' law of gravity

$$\nabla \cdot \mathbf{g} = -4\pi G\rho$$

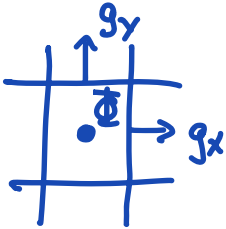
where G is gravitational constant

Combine to obtain Poisson's eqn for gravity

$$\nabla^2\Phi = 4\pi G\rho$$

Similar to groundwater: $\mathbf{g} \rightarrow \mathbf{q}$
 $\Phi \rightarrow h$

If we discretize



$$\underline{g} = \begin{bmatrix} g_x \\ g_y \end{bmatrix}$$

$$\underline{D} * \underline{G} * \underline{\Phi} = 4\pi G \rho$$

$$\underline{L} \underline{\Phi} = \underline{f}_s$$

$$\underline{g} = - \underline{G} * \underline{\Phi}$$

⇒ \underline{g} is like a flux and lives on faces

Gravity term in Darcy's law

$$\underline{q} = - \frac{k}{\mu} (\nabla p + \rho_f \underset{\substack{\uparrow \\ \text{scalar}}}{g} \hat{z})$$

$$g = |\underline{g}|$$

Reformulate Darcy's law in terms of g

$$\underline{q} = - \frac{k}{\mu_f} (\nabla p + \rho_f \underset{\substack{\uparrow \\ \text{scalar}}}{g} \underset{\substack{\uparrow \\ \text{vector}}}{\hat{z}})$$

$$\hat{z} = \nabla z$$

introduce:

$$\underline{g} = - g \nabla z$$

↑ vector ↑ scalar

Note: minus sign because g points downwards

Darcy's law in terms of gravitational field vector

$$\mathbf{q} = -\frac{k}{\mu_f} (\nabla p - \rho_f \mathbf{g})$$

substituting into the mass balance

$$\nabla \cdot \mathbf{q} = f_s$$

we have the following Poisson eqn. for pressure

$$-\nabla \cdot \left[\frac{k}{\mu_f} \nabla p \right] = f_s - \nabla \cdot \left[\frac{k}{\mu_f} \rho_f \mathbf{g} \right]$$

k = intrinsic permeability $\rightarrow x_c$

μ_f = dynamic viscosity of fluid

here we assume it is constant, but generally

function of $T \rightarrow x_c$

ρ_f = density of fluid, depends on $T \rightarrow x_c$

\mathbf{g} = gravitational field vector $\rightarrow x_f$

$\lambda = \frac{k}{\mu_f}$ = mobility $\rightarrow x_c$

Discretization of Lhs is the same

$$-\nabla \cdot [\lambda \nabla p] \approx -\underline{D} * \underset{\substack{\uparrow \\ \text{haru. average of } \lambda's}}{\underline{Kd}} * \underline{G} * p = \underline{L} * p$$

Discretize rhs:

$$f_s - \underbrace{\nabla \cdot \left[\frac{k}{\mu_f} \rho_f g \right]}_{f_g} \approx \underline{f}_s + \underline{f}_g$$

⇒ new rhs. vector \underline{f}_g due to gravity

$$\text{New variables: } \left. \begin{array}{l} \text{grav} = g \\ \text{grav_vec} = \underline{g} \end{array} \right\} g = |\underline{g}|$$

$$\underline{f}_g = -\underline{D} * \underline{Kd} * \underline{Rho} * \underline{\text{grav_vec}}$$

$$\underline{\text{grav_vec}} = \underline{G} * \text{Grid.xc} \quad \text{in 1D}$$

$$= \underline{G} * Y_c(:) \quad \text{in 2D}$$

where Y_c is generated by meshgrid

$$\underline{Lmu} = \text{comp_mean}(\underline{Lmu}, \underline{M}, -1, \text{Grid}, 1);$$

$$\underline{Rho} = \text{comp_mean}(\underline{rho}, \underline{M}, 1, \text{Grid}, 1);$$

Note: • It is not entirely obvious how p on faces should be estimated

- f_g is zero on boundaries